

libRadtran user course, lecture # 8

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The libRadtran radiative transfer package

The libRadtran software package is a suite of tools for radiative transfer calculations in the Earth's atmosphere.

- Borne sometime in the early 90s.
- Early development A. Kylling and B. Mayer.
- First publication with uvspec in the title Mayer et al. (1997).
- Continuous development until today.
- Freely available from www.libradtran.org.
- Currently version 2.0.2. See <http://www.geosci-model-dev-discuss.net/8/10237/2015/gmdd-8-10237-2015.html>.
- More than 600 papers have actively used libRadtran.

libRadtran: some applications

- Analysis of **UV-radiation** measurements
- **Cloud and aerosol remote sensing** using solar and thermal measurements.
- **Volcanic ash studies**
- **Remote sensing of surface properties**
- **Trace gas remote sensing**: forward modelling
- Calculation of **actinic fluxes**
- Determination of solar direct irradiance and global irradiance distributions.
- Simulation of satellite radiances to be used for data assimilation
- Validation of radiation schemes included in **climate models**
- Simulation of heating rates in 3D
- Simulation of solar radiation during a **total eclipse**
- Rotational **Raman scattering**, which explains the filling-in of Fraunhofer lines
- Remote sensing of **planetary atmospheres**

libRadtran tools

Some libRadtran tools are:

uvspec The main radiative transfer tool

mie Calculation of single scattering properties

ssradar Single scattering radar simulator

zenith calculates solar zenith and azimuth angles and Earth eccentricity

angres calculates the effective radiation given an angular response and a radiance distribution

vac2air converts from vacuum to air wavelength and vice versa

snowalbedo calculates diffuse and direct albedo of snow

make_slitfunction generates various slit functions

conv convolves an input file with a convolution function

integrate integrates a file between given limits

spline spline or linear interpolation between given data points

More are included in the package.

Useful folders

- bin** Contains all executeables including `uvspec`.
- doc** The libRadtran User's guide is found here.
- examples** Look here for examples of input and, the resulting output, files.
- GUI** Location of the Graphical User Interface to `uvspec`.
- src** The source code for the main tools.
- libsrc_c** C source that is used by many tools.
- libsrc_f** Fortran source that is used by many tools.
- src_py** Place to start if you want to add your own input option (python).

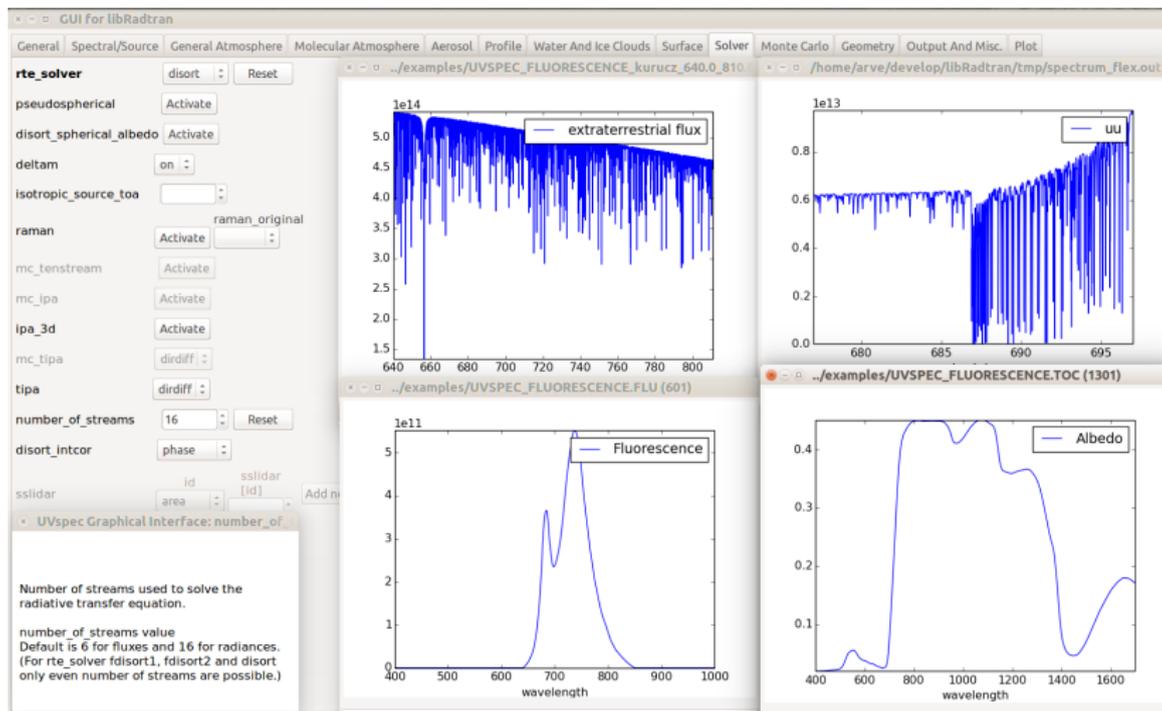
Calculate solar zenith angle:

```
zenith -q -a 60 -o -20 -e -s -15 19 1 12 00  
12:00:00 60.6561 3.4746 1.009064
```

Add following to input file

```
albedo 0.2 # Surface albedo  
sza 60.6561 # Solar zenith angle  
mol_modify O3 300. DU # Set ozone column  
day_of_year 19 # Correct for Earth-Sun distance  
rte_solver disort # Radiative transfer equation solver
```

Graphical User Interface (GUI)



Repeated runs cumbersome. Use scripting languages to generate input file, run code and handle output:

`python` Powerful and easy to use.

`shell` Various unix shells

`perl` Powerful scripting language. If you find shell scripting syntax awkward you do not want to try perl

GUI simple visualization of some input/output

gnuplot useful for quick views of data in column format

python/matplotlib Can do nearly everything

python/mayavi Powerful 3D plotting

Ozone retrieval (look-up tables)

Use Perl script: `src/Gen_o3_tab.pl`. **WARNING: options not updated to version 2.0.**

`Gen_o3_tab` generates a table of some wavelength ratio versus solar zenith angles for different ozone amounts. The table is read and ozone columns derived by the `read_o3_tab` program.

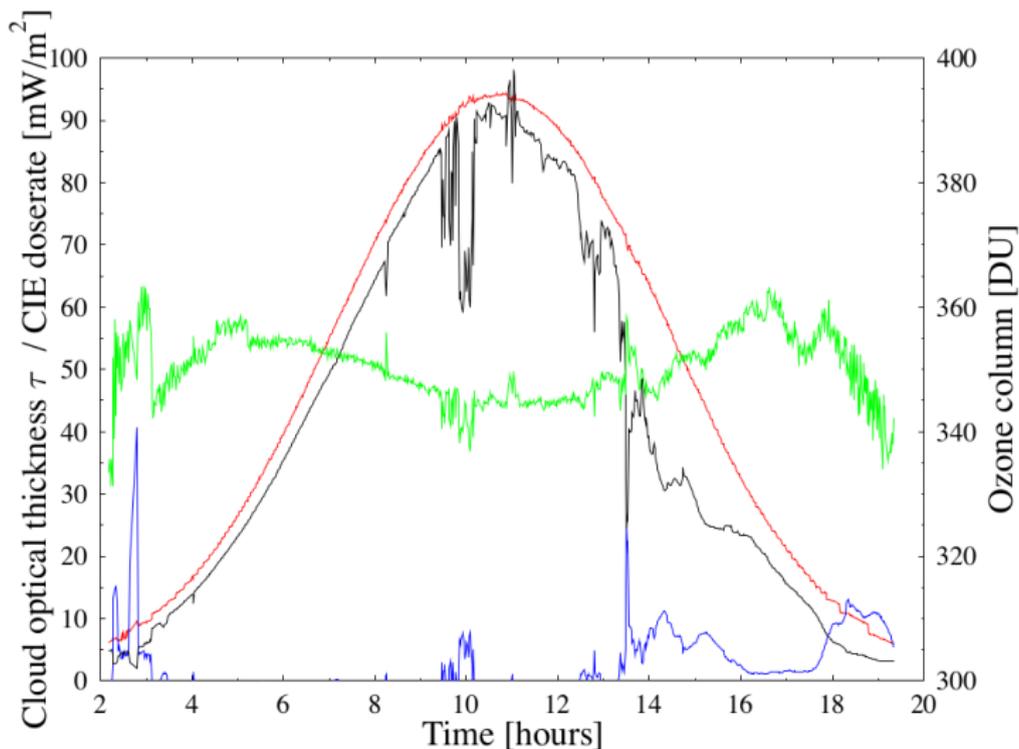
The following wavelength ratios are supported:
(See the `libRadtran` documentation for examples.)

- * Ratios between two wavelengths where each is calculated from single wavelength measurements.
- * Ratios between two wavelengths where each is calculated from wavelength measurements multiplied by a bandpass function and integrated over the bandpass.

`Gen_o3_tab` understands the following options:

```
--absolute           : The wavelengths of the bandpass files are in
                       : absolute units. Default is relative units.
--albedo <value>    : Lambertian surface albedo. Default is 0.0.
--alpha <value>     : Angstrom alpha coefficient. Default is 0.0.
--altitude <value>  : Altitude above sea level [km]. Default is 0.0.
--atmmod <name>     : Name of atmosphere file. Default atmmod/afglus.dat.
--beta <value>      : Angstrom beta coefficient. Default is 0.0.
--help              : Prints this message.
--o3_crs <name>     : Name of o3 cross section to use. Default is Molina.
                       : See libRadtran documentation for other options.
--slitfunction <name> : Name of slitfunction file.
--bandpasslower <name> : Name of file holding bandpass for lower wavelength.
--bandpassupper <name> : Name of file holding bandpass for upper wavelength.
--file <name>       : Name of file where the table will be stored.
--lower_lambda <value> : Value for lower wavelength, in nm.
--upper_lambda <value> : Value for upper wavelength, in nm.
--zenith            : Calculate zenith sky radiance table.
```

Ozone retrieval cont'd



From Mayer and Kylling (2005).

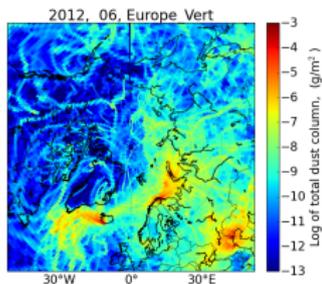
Global radiation (SW+LW)

- 1 Compare "accurate" radiative transfer with met & climate models.
- 2 Radiative forcing (aerosol direct effect)
 - Calculate irradiances with and without forcing (aerosol).
 - Include water and ice clouds
 - Aerosol from dispersion model
 - Assume independent pixels
 - Calculate solar and thermal radiation separately
 - Computationally expensive if done for many pixels and long time periods.

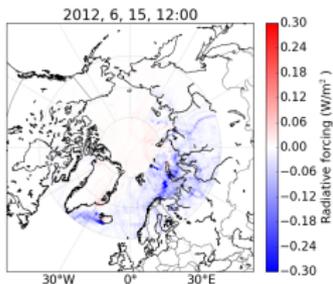
Useful options:

```
wc_file 1D filename
ic_file 1D filename
output_user .....
mol_file o3, co2, h2o .....
zout ....
output_process sum
pseudospherical
cloud_fraction_file .....
mol_abs_param fu
rte_solver disort
day_of_year ...
albedo_file ....
sur_temperature ....
sza ....
profile_file ...
profile_properties ...
```

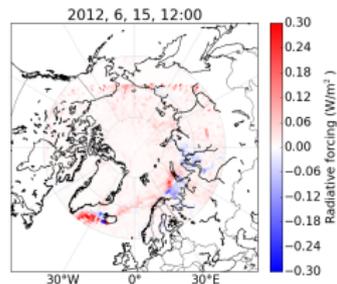
Radiative forcing (aerosol direct effect)



Dust column



Cloudless



Cloudy

RF simulations only made north of 60°N .

Photolysis rates (aerosol)

Photolysis rate for reaction i given by

$$J_i(\tau) = 4\pi \int d\lambda q_i(\lambda) \sigma_i(\lambda) \overline{I(\tau, \lambda)}$$

where the mean intensity

$$\overline{I(\tau)} = \frac{1}{2\pi} \left[I_0 e^{-\tau/\mu_0} + \int_0^{2\pi} d\phi \int_0^1 I(\tau, -\mu, \phi) d\mu + \int_0^{2\pi} d\phi \int_0^1 I(\tau, \mu, \phi) d\mu \right]$$

uvspec may calculate the mean intensity.

`src/Calc_J.pl` may calculate some photolysis rates. **WARNING: options not updated to version 2.0.**

- Latest version always at `www.libradtran.org`.
- Missing something for your research?
 - Code is available so you may add yourself. If so please send modified code to us so we can include in future releases.
 - We may include it, but funding is needed.
- We are always interested in joining projects related to radiative transfer (modelling, measurement, data interpretation and analysis).

References I

- Mayer, B. and Kylling, A.: Technical note: the libRadtran software package for radiative transfer calculations-description and examples of use, *Atmos. Chem. Phys.*, 5, 1855–1877, 2005.
- Mayer, B., Seckmeyer, G., and Kylling, A.: Systematic long-term comparison of spectral UV measurements and UVSPEC modeling results, *J. Geophys. Res.*, 102, 8755–8767, 1997.